

# Danish 1.54-m Telescope, La Silla

## Allocation schedule, Danish and Chilean time

Periods 71 + 72, March 2003 – March 2004

Instrument: DFOSC

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Dates, from noon to noon	P.I.	Programme
23.03 – 26.03		Technical time
26.03 – 02.04 <sup>1)</sup>	R. Michelsen	Spectroscopy of Near-Earth Asteroids
02.04 – 11.04 <sup>1)</sup>	M. I. Andersen	The mass of the Milky Way
11.04 – 20.04 <sup>1)</sup>	B. Nordström	BVRI photometry of extremely metal-poor stars
20.04 – 21.04	J. Hjorth	Observations for educational purposes
23.06 – 27.06	E. Costa	Photometry for nearby southern stars
27.06 – 04.07	M. Catelan	Sagittarius DSPH globular clusters
04.07 – 10.08 <sup>2)</sup>	U. G. Jørgensen	Microlensing search for extra-solar Jovians
10.08 – 18.08 <sup>1),3)</sup>	T. Arentoft	Mapping of stars around MONS targets
16.09 – 19.09		Technical time
19.09 – 30.09 <sup>1)</sup>	L. F. Olsen	$z = 0.3 - 0.4$ EIS cluster candidates
30.09 – 08.10 <sup>1)</sup>	J. Pritchard	Magellanic Cloud eclipsing binaries
08.10 – 16.10 <sup>1)</sup>	B. Nordström	BVRI photometry of extremely metal-poor stars
15.11 – 19.11	E. Costa	Photometry for nearby southern stars
19.11 – 02.12 <sup>1)</sup>	J. Pritchard	Magellanic Cloud eclipsing binaries
02.12 – 14.12 <sup>1)</sup>	T. Dall	Mode characterization in $\delta$ Scuti stars
13.01 – 17.01 <sup>1)</sup>	D. Stello	Oscillations in M67 giant stars
17.01 – 24.01 <sup>1)</sup>	L. F. Olsen	$z \cong 0.2$ EIS cluster candidates
24.01 – 11.02 <sup>1)</sup>	D. Stello	Oscillations in M67 giant stars
11.03 – 15.03		Technical time
15.03 – 19.03	E. Costa	Photometry for nearby southern stars
19.03 – 01.04 <sup>1)</sup>	T. Dall	Mode characterization in $\delta$ Scuti stars

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1)

The observer is expected to collaborate on the ToO programme “ $\gamma$ -ray bursts afterglows and host galaxies” by J. Hjorth et al. The target of opportunity observations will consist of  $\sim 1$  hour for early identification and max. 5 hours for UBVR follow-up imaging. The ToO programme is limited to max. 6 hours per observing run, max. 2 events in period 71, and max. 6 events in period 72.

2)

20% of the time in this long period is allotted to the programme “ $\gamma$ -ray bursts afterglows and host galaxies” by J. Hjorth et al.

3)

If the Rømer/Mons project is cancelled, this period is allotted to the programme “Mode characterization in  $\delta$  Scuti stars” by T. Dall et al.

Poul E. Nissen